

UCI PHYSICS/CHEM207 – Applied Physical Chemistry, Summer 2022

Lecture #15 of 14

(3: TThF, 5: MTWThF, 4: MTWTh, 2: TW<u>Th</u>) {Okay... let's wrap this up}

> Prof. Shane Ardo Department of Chemistry University of California Irvine



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Photochemistry

Prof. Shane Ardo Department of Chemistry University of California Irvine

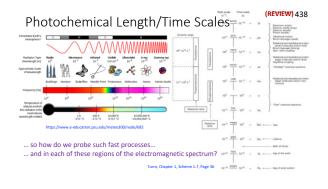
Photochemistry

437

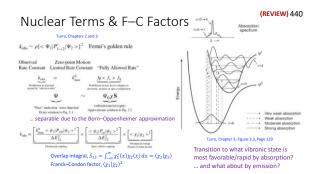
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436

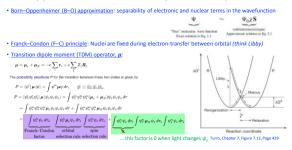
- Blackbody radiation, Carnot efficiency limits, Light–Matter interactions, Photon properties, Conservation laws
- Jablonski diagram, Internal conversion, Intersystem crossing, Kasha– Vavilov rule, Thexi state, Stokes shift, Luminescence processes
- Harmonic oscillator model, Born–Oppenheimer approximation, Franck–Condon principle, Transition dipole moment operator, Selection rules, Spin–orbit coupling, Heavy-atom effect
- Photochemical length and time scales, Electromagnetic spectrum
- Beer–Lambert law, Absorption coefficient, Einstein coefficients, Oscillator strength, Absorptance, *E*–*k* diagrams



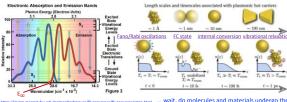
(REVIEW) 439 Electromagnetic spectrum ... what can one do with microwaves? rotational 10nm 1nm 1µm MIR NIR 0.1 FIR vův xuv мw w. vibrational levels, W_y" ... and what about gamma rays ... and neutrons... and electrons



B–O Approximation, F–C Principle, TDM Operator



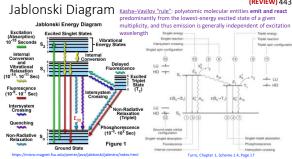
Thermally Equilibrated Excited (Thexi) State



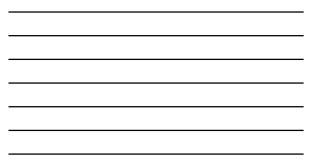
... and why are these spectra plotted as a function of wavenumber... and not wavelength? ... so that you can see the <u>mirror-image "rule"</u>

html ... wait, do molecules and materials undergo the same physical processes and follow the same laws of the Universe?... shocking, isn't it?!?!?!

P. Narang, R. Sundararaman & H. A. Atwater, Nanophoton., 2016, 5, 96–111



(REVIEW) 443



Selection Rules

Angular Momentum Quantum Numbers (REVIEW) 444

 $\underbrace{\int \psi_{\tau}^{*} \psi_{v} \, d\tau_{n}}_{\text{Franck-Condon}} \underbrace{\int \psi_{c}^{*} \mu_{c} \psi_{c} \, d\tau_{c}}_{\text{orbital}} \underbrace{\int \psi_{s}^{*} \psi_{s} \, d\tau_{s}}_{\text{spin}}$

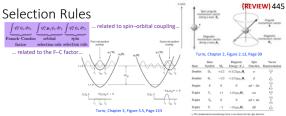
Photon... which came from matter: s = 1, $m_s = \pm 1$ Electron (Orbital): l, $m_l = [-l, l]$ in steps of 1 Electron (Spin): $s = \frac{1}{2}$, $m_s = \left[-\frac{1}{2}, \frac{1}{2}\right]$

... well these are just overlaps... and so the more overlap, the more favorable a transition... ... the F–C (nuclear vibrational) factor makes sense based on pictures on the previous slides ... but what does μ_c do to a wavefunction?... it uses it to make it coincide with an unoccupied orbital...

... but what does μ_e do to a wavefunction?... it uses \overline{E} to make it coincide with an unoccupied orbital. and even if we didn't know, it better change the angular momentum during photon annihilation ... and what are spin wavefunctions?... just symbols!... spin does not fall out of $\mu_{\rm ch}$ it's just math...

so, the spin wavefunctions only overlap when they are identical... meaning spin does not change

... simplified systems... (Angular Momentum) Atomic Selection "rules" Orbital angular momentum (Laporte "rule"): $\Delta l = \pm 1...$ as $l_f = l_t \pm s_{\rm photon}$ Spin angular momentum (Wigner "rule"): $\Delta m_s = 0... \mu_e$ does not act on spin Orbital z-direction angular momentum: $\Delta m_l = 0, \pm 1...$ as $m_{l,f} = m_{l,l} \pm m_{\rm sphoton}$... the allowed 0 option can be envisioned as two vectors that are opposite in one direction



... simplified systems... Summary of Atomic Selection "rules" Turns, Obspire 2, Table 24, Pige 102 $\Delta I = \pm 1$, since $I_f = I_{\downarrow} \pm s_{\rm photon}$. $\Delta m_s = 0$... $\Delta m_t = 0, \pm 1$, since $m_{t,f} = m_{t,1} \pm m_{s,\rm photon}$. Heavy Molecule (Russell–Saunders L–S Coupling) Selection "rules"... for linear oscillating photon \vec{E} . Total angular momentum: $\Delta f = 0, \pm 1$... and $\Delta S = 0$... and $\Delta L = 0, \pm 1$.

Total z-direction angular momentum: $\Delta m_{J}=0,\pm1...$ and 0's are there for the reason before

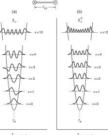
Selection Rules

 $\underbrace{\int \psi_c^* \psi_c \, d\tau_c}_{\mathbf{X}} \int \psi_c^* \boldsymbol{\mu}_N \psi_v \, d\tau_v \int \psi_s^* \psi_s \, d\tau_s$

When light does not change $\psi_{e^{,...}}$... this factor is non-zero... and the other factor is 0 This means that the photon absorption event... ... is nuclear... and is not electronic

 $\label{eq:summary of Nuclear Selection "rules"} \\ \mbox{Vibrational (Harmonic Oscillator):} \\ \Delta v = \pm 1 \mbox{ (change in dipole)} \\ \mbox{Vibrational (Harmonic Oscillator) <u>Scattering:</u>} \\ \end{tabular}$

 $\Delta v = \pm 1$ (polarizable) Rotational (Rigid Rotor Spherical Harmonics): $\Delta J = \pm 1$ (permanent dipole) (REVIEW) 446



Turro, Chapter 2, Figure 2.6, Page 76

447 Absorption Coefficient & Beer–Lambert Law To describe attenuation of light intensity/power through matter due to absorption only... one writes $\frac{\partial I_v}{\partial z} = -\alpha_v I_{v}$, where α_v is the <u>linear Napierian</u> absorption coefficient (cm⁻¹) σz $\ln \left(\frac{I_{\nu,\text{back}}}{I_{\nu,\text{front}}}\right) = -\alpha_{\nu}\ell_{\cdots} \text{ or } I_{\nu} = I_{\nu,0}e^{-\alpha_{\nu}\ell}, \text{ where } I_{\nu} = I_{\nu,\text{back}} \text{ and } I_{\nu,0} = I_{\nu,\text{front}}$ (vy,ront/ ... where $T_{\nu} = \frac{l_{\nu}}{l_{\nu,o}}$ (transmittance) and $A_{\nu} = -\log(T_{\nu}) = \log\left(\frac{l_{\nu,o}}{l_{\nu}}\right)$ (absorbance) ... but the **absorption coefficient** can take on many forms/units... sorry...

 $\log \left(\frac{l_v}{l_{v,o}}\right) = -a_v \ell_{v,v}$ where a_v is the linear decadic absorption coefficient (cm⁻¹) [not often used] $\ln\left(\frac{i_{\nu}}{i_{\nu,o}}\right) = -\kappa_{\nu}c\ell...$ where κ_{ν} is the <u>molar Napierian</u> absorption coefficient (M⁻¹ cm⁻¹) [n. o. u.] ... since M⁻¹ cm⁻¹ = L mol⁻¹ cm⁻¹ = dm³ mol⁻¹ cm⁻¹, $\sigma_{\nu} = \frac{1000 \kappa_{\nu}}{N_{\rm A}}$ is the **absorption** cross-section (cm²)

 $\log\left(\frac{I_{\nu}}{I_{\nu\sigma}}\right) = -\varepsilon_{\nu}c\ell...$ where ε_{ν} is the <u>molar decadic</u> **absorption coefficient** (M⁻¹ cm⁻¹)... finally!...

... leading to the **Beer–Lambert law**... $A_{\nu} = \varepsilon_{\nu} c \ell$... a succinct and well-known equation in the end

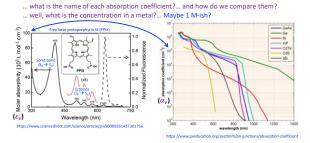
448 Einstein Coefficients & Oscillator Strength

<u>Oscillator strength</u> (f_{12}): integrated strength of an absorption band relative to a completely ... is emission the microscopically reversible opposite of absorption? ... Nopel allowed transition Absorption = -Stimulated Emission... but not Spontaneous Emission Spontaneous Emission Positive Absorption = Stimulated Absorption ntion Stin $\frac{\partial n_1}{\partial n_1} = A_{21}n_2$ $\frac{\partial n_1}{\partial t} = -B_{12}n_1\rho(\nu)$ Stimulated abso lated emission Negative Absorption = Stimulated Emission $\frac{\partial n_1}{\partial v} = B_{21} n_2 \rho(v)$ ∂t Consider all possible reactions to be at equilibrium... Figur $\frac{\partial a_1}{\partial t} = -\frac{\partial a_2}{\partial t} = \mathbf{0} = A_{21}n_2 + B_{21}n_2\rho(\nu) - B_{12}n_1\rho(\nu) \dots g_1B_{12} = g_2B_{21} = \frac{g_1c\sigma_{12}}{h\nu} = \frac{g_1c\sigma_{12}}{4\varepsilon_0m_0h\nu}$ hv

 $\frac{\partial t}{\partial t} = \frac{\partial t}{\partial t} =$

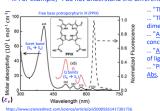
Absorption Coefficient

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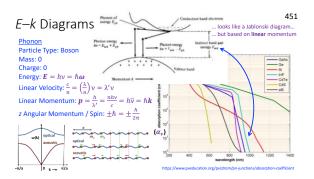
Absorption vs Absorbance vs Absorptance

... let's use each in a sentence, okay?



... For example, "I do not understand the difference between A., A., and A.?"... Ha! j/k "This plot displays an absorption spectrum" ... "The ordinate is often <u>absorbance</u>, which is dimensionless... and definitely not "a.u."" "Absorbance, A, is directly related to concentration... and so its range is $[0, \infty)$ " ... "<u>Absorptance</u>, α , is the fraction (or percentage) of light absorbed... it goes well with T (or %T)" ... "Photochemists must know concentrations (via Abs, A) and amount of light absorbed (as Abt, α)' $\% R_{\nu} = \left(\frac{n_2 - n_1}{n_2 + n_1} \right)^2$

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Photochemistry (summary for today)

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Photochemistry (*let's review* TWTh)

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- W: Jablonski diagram, Internal conversion, Intersystem crossing, Kasha-Vavilov rule, Thexi state, Stokes shift, Luminescence processes, Harmonic oscillator model, Born–Oppenheimer approximation, Franck– Condon principle, Transition dipole moment operator, Selection rules, Spin–orbit coupling, Heavy-atom effect, Photochemical length and time scales, Electromagnetic spectrum
- Th: Beer–Lambert law, Absorption coefficient, Einstein coefficients, Oscillator strength, Absorptance, E-k diagrams

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ALL -SO MUCH FOR TAKING PART II, (RATHER) NEW COURSE! ... THANKS SO MUCH FOR TAKING PART IN THIS

(1) Some chemists know a lot of physics; (2) Physicists (you!) now know a lot of Chemistry;

(3) Who wants to switch to a Ph.D. in Chemistry?