

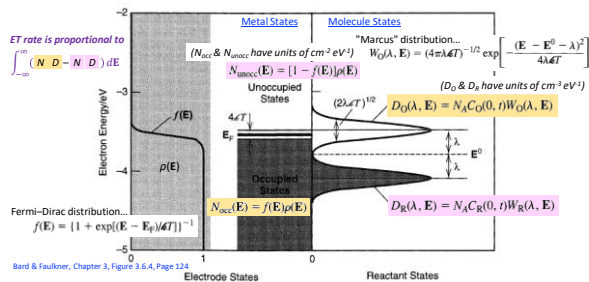


Lecture #8 of 12

Prof. Shane Ardo
 Department of Chemistry
 University of California Irvine

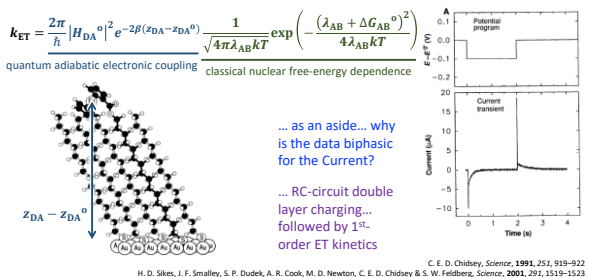
Marcus–Gerischer Theory

(REVIEW) 185



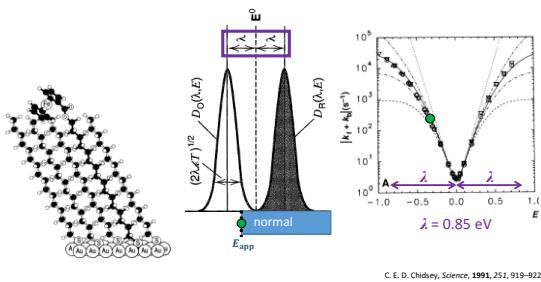
Marcus–Gerischer Theory

(REVIEW) 186



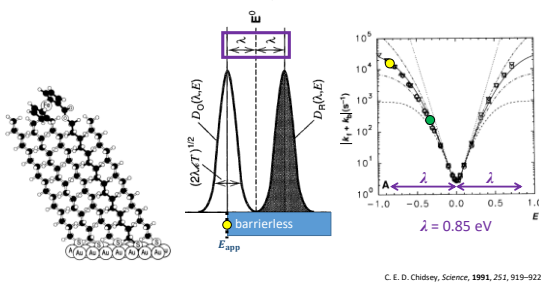
Marcus–Gerischer Theory

(REVIEW) 187



Marcus–Gerischer Theory

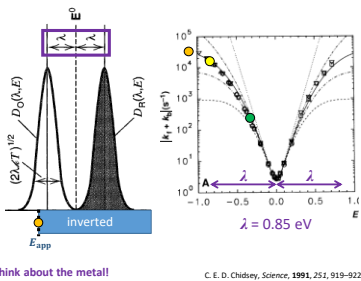
(REVIEW) 188



Marcus–Gerischer Theory

(REVIEW) 189

- It is easy to sweep/vary the driving force, ΔG_{app} , by simply changing the electrochemical potential of electrons (e^-) in the [Metal] working electrode, $\mu_{e^-}^{M}$, through variations in E_{app}
 - But evidence of the inverted region is a little challenging to clearly observe
- ... what if Chidsey had plotted the derivative of his data on the right?
- ... what do you expect that would have looked like?
- ... a nice Marcus parabola!
- ... I wish he had done that!
- ... anyway... why is λ so small in water? ... think about the metal!

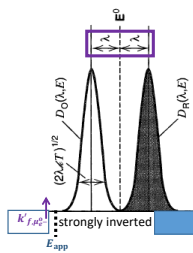


Marcus–Gerischer Theory

190

• Use of a semiconductor limits the electronic states to those with (approximately) a single μ_e° , which makes analysis of data simpler, i.e. one does not need to consider a distribution of states in the electrode

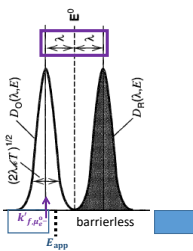
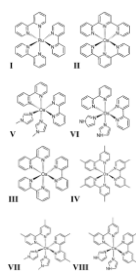
• But one cannot alter the driving force, ΔG_{app}° , by simply changing the electrochemical potential of electrons (e^-) in the (S)emi(C)onductor working electrode, $\bar{\mu}_e^{SC}$, through variations in E_{app} , because instead that changes the concentration of e^-



How can one use a semiconductor to study the inverted region?
Think solution studies... vary the molecule!

Marcus–Gerischer Theory

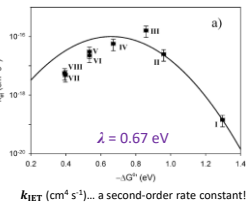
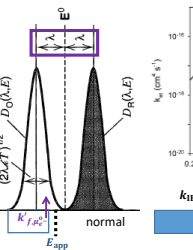
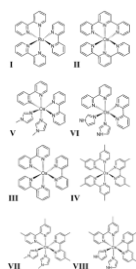
191



T. W. Hamann, F. Gotrein, B. S. Brunenschwig & N. S. Lewis, *J. Am. Chem. Soc.*, **2005**, *127*, 7815–7824
T. W. Hamann, F. Gotrein, B. S. Brunenschwig & N. S. Lewis, *J. Am. Chem. Soc.*, **2005**, *127*, 13949–13954

Marcus–Gerischer Theory

192



T. W. Hamann, F. Gotrein, B. S. Brunenschwig & N. S. Lewis, *J. Am. Chem. Soc.*, **2005**, *127*, 7815–7824
T. W. Hamann, F. Gotrein, B. S. Brunenschwig & N. S. Lewis, *J. Am. Chem. Soc.*, **2005**, *127*, 13949–13954

k_{ET} ($cm^2 s^{-1}$)... a second-order rate constant!



Photophysical Processes

Prof. Shane Ardo
Department of Chemistry
University of California Irvine

Today's Critical Guiding Question

194

What continuity/conservation laws are most important for photophysical processes like absorption and emission of photons... for real this time: Part 2?

Photophysical Processes

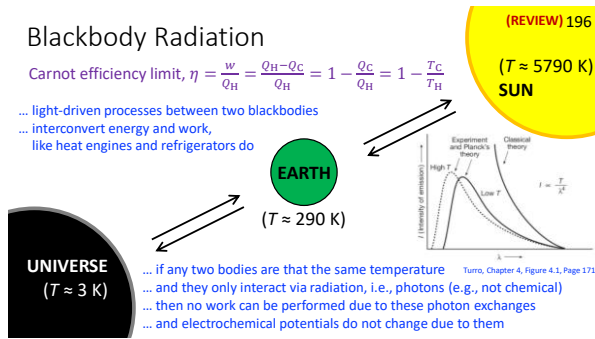
(UPDATED) 195

- Blackbody radiation, Photon properties, Light–Matter interactions, Conservation laws, Einstein coefficients
- Jablonski diagram, Spin multiplicity, Internal conversion, Intersystem crossing, Thexi state, Kasha–Vavilov rule, Stokes shift, PL
- Born–Oppenheimer approximation, Franck–Condon principle, Transition dipole moment operator, Franck–Condon factors, Beer–Lambert law, Absorption coefficient, Oscillator strength, Absorptance
- Luminescence processes, Selection rules, Charge-transfer transitions, Spin–Orbit coupling, Heavy-atom effect, $E-k$ diagrams, Jortner energy gap law, Conical intersections, Energy transfer, Exciplex/Excimer
- Photoluminescence spectrometer, Emission/Excitation spectra, Inner filter effects, Anisotropy, Excited-state lifetime, Emission quantum yield

Blackbody Radiation

Carnot efficiency limit, $\eta = \frac{W}{Q_H} = \frac{Q_H - Q_C}{Q_H} = 1 - \frac{Q_C}{Q_H} = 1 - \frac{T_C}{T_H}$

... light-driven processes between two blackbodies
 ... interconvert energy and work,
 like heat engines and refrigerators do



... if any two bodies are that the same temperature
 ... and they only interact via radiation, i.e., photons (e.g., not chemical)
 ... then no work can be performed due to these photon exchanges
 ... and electrochemical potentials do not change due to them

Photon Properties & Conservation Laws

(REVIEW) 197

Where does light come from?

Particle Type: Boson

Mass: 0

Charge: 0

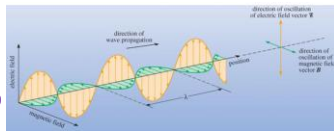
Energy: $E = h\nu = \hbar\omega$

Linear Velocity: $\frac{c}{n} = \left(\frac{\lambda}{n}\right) \nu = \lambda' \nu$

Linear Momentum: $p = \frac{h}{\lambda'} = \frac{\hbar\nu}{c} \approx 0$

Linear Polarization: \vec{E} and \vec{B}

z-Direction Angular Momentum / Circular Polarization / Chirality / Helicity / Spin: $\pm \hbar = \pm \frac{h}{2\pi}$



Wait... is a light a wave or a particle?
 ... I mean, is matter a wave or particle?
 ... I mean, doesn't everything exhibit
 wave-like and particle-like properties?

With what matter does light interact?

Fermion Angular Momentum (Orbital, Spin)

Magnitude: $\hbar\sqrt{j(j+1)}$

z-Direction: $m_j \hbar, m_j = [-j, j]$ in steps of 1

Multiplicity/Degeneracy, $g_j: 2j + 1$

Light-Matter Interactions

198

Total force exerted on an electron by a light wave

$$\mathbf{F} = e\mathbf{E} + \frac{e[\mathbf{H}\nu]}{c}$$

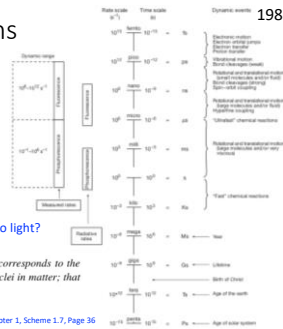
Electrical force Magnetic force

$c_{light} = 3 \times 10^{18} \text{ nm s}^{-1}$

$v_{electron} = 10^{15} - 10^{16} \text{ nm s}^{-1}$

$v_{nuclei} = 10^{13} - 10^{14} \text{ nm s}^{-1}$

... so which term dominates the resonant response to light?

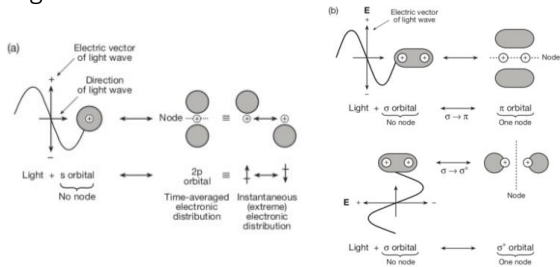


In more concrete chemical terms, the oscillation of the dipoles corresponds to the movements of electrons in bonds relative to positively charged nuclei in matter; that is, electrons oscillate about the nuclear framework of molecules.

Turro, Chapter 1, Scheme 1.7, Page 36

Light-Matter Interactions

199



Turro, Chapter 4, Figure 4.6, Page 189

Light-Matter Interactions

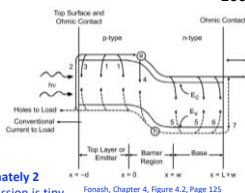
200

$$\frac{\partial c_{A,z_0}}{\partial t} = \sum_j R_{A,j} - \frac{\partial N_A}{\partial z}$$

What value of j have we considered thus far? ≥ 2

How large is j for actual systems? **Quite large, likely!**

What is the smallest value that j can be? **3... but approximately 2**
 ... stimulated emission is tiny



Fonash, Chapter 4, Figure 4.2, Page 125

Given a box at temperature, T , by what processes can heat be transferred to something inside it?
 Okay, now what if inside the box was a vacuum?

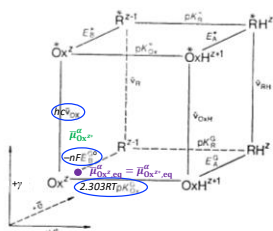


... at a microscopically reversible equilibrium, rate is equal to % $\lambda_A(v)$ x PhotonFlux(v), integrated over v
 ... $\bar{\mu}_A = \bar{\mu}_{A^*}$... with additional (sun)light absorption, $\bar{\mu}_A < \bar{\mu}_{A,eq}$ and $\bar{\mu}_{A^*} > \bar{\mu}_{A^*,eq}$ = useful work!

... okay... now let's try this again...

(REVIEW) 201

Förster Cube and Square Schemes



... all of these free energy terms are standard-state free energies (ΔG°)... but what is the actual free energy of the system (ΔG)?

... let's assume that $\Delta G = 0$ (equilibrium)... how could I indicate that on this slide, as a point(s), to depict the majority species present?

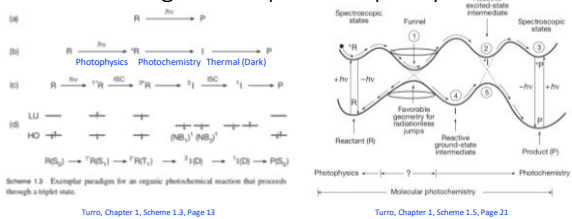
... now, how can one push/pull this system out of equilibrium?

... recall Le Chätelier's principle... and thus by addition of reactants or removal of products... such as mass or light!

... hopefully this made a little more sense this time around... and if not, let's keep on trying!

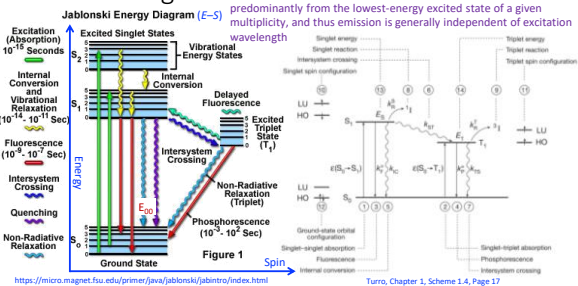
Z. R. Grabowski & W. Rubaszewska, J. Chem. Soc. Faraday Trans. 1, 1977, 73, 11-28

Jablonski Diagram & Spin Multiplicity 202



What is the origin of the names "singlet" and "triplet"? ... Angular Momentum Energy Degeneracy, $g_j = 2J + 1$
 ... when $J = 0, g_j = 1$... sounds like a "Singlet (S or :X)"
 ... when $J = 1, g_j = 3$... sounds like a "Triplet (T or :X)"

Jablonski Diagram 203



Today's Critical Guiding Question 204

What continuity/conservation laws are most important for photophysical processes like absorption and emission of photons... for real this time: Part 2?
